

## Pcal at the AuScope antennas

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10th IVS General meeting 2018, June 3-8, Svalbard IVS Analysis WS

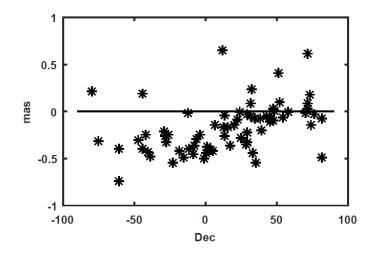
**Declination bias in CRF solution** 

Differences between manual pcal and normal pcal sessions found (EVGA 2017)



#### **Testbed Cont14**

# Reprocessing of the Cont14 files repeat fringe fitting



XA solution (applying pcal), declination bias clearly visible)

MX ... applying manual pcal in Hb

→ Declination bias is still clear, though a bit mitigated

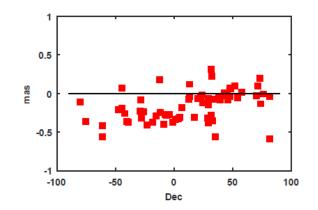


Fig. 2 Estimated source declination with respect to ICRF2 as determined in a global solution of the XM sessions. Here, manual pcal was applied at Hb. At the moment, this solution comprises 7 sessions, c1401 through c1407. The declination bias is still clearly visible, though a bit mitigated.

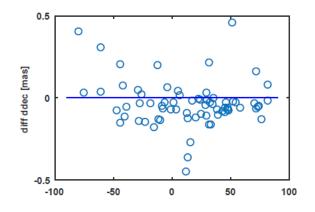


Fig. 3 Difference between the estimated source declination with respect to ICRF2 between the XA and the XM solutions. Negative values mean the declination bias becomes smaller in the XM solution.



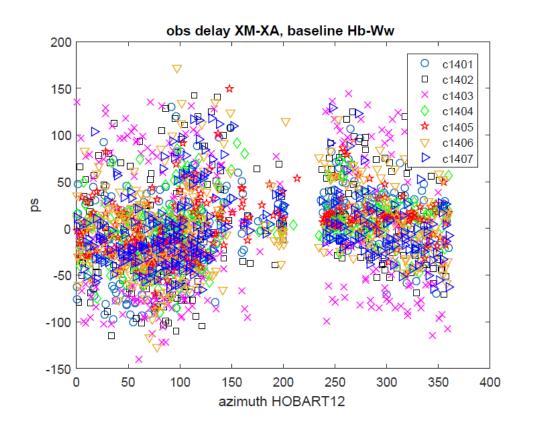


Fig. 5 Difference in the observed delays between the XM and the XA sessions. The comparison was done per baseline, in this case on the Hb-Ww baseline. The differences are color-coded per sessions and a mean value per session was removed.



### Systematic delay

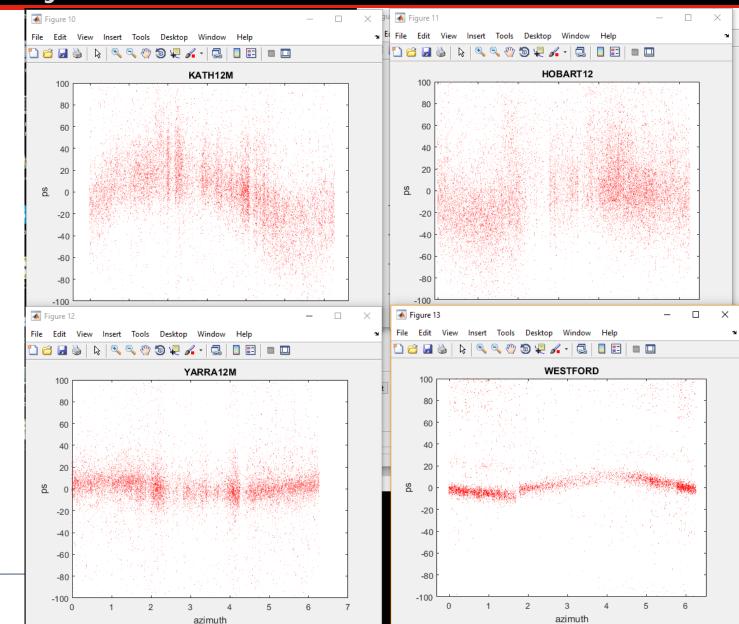
Pcal directly calculated from vgosDB phascal phases

All sessions in 2014

Long-term temperature effect removed

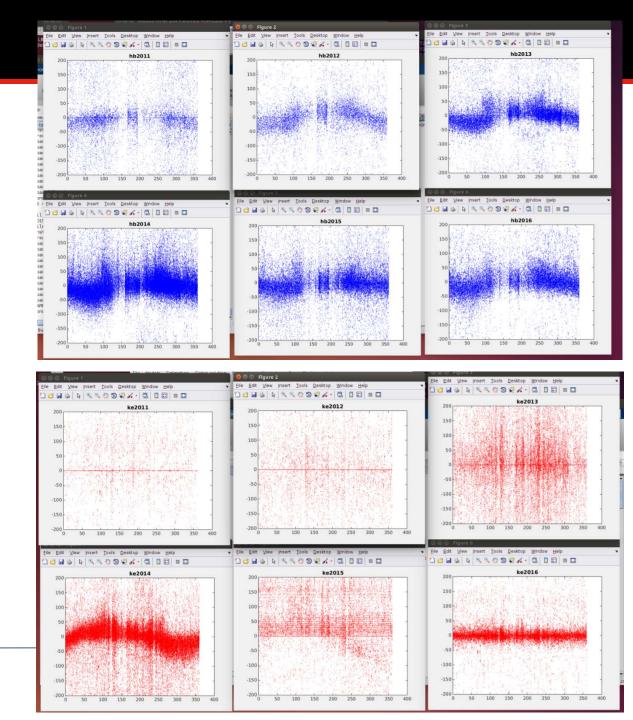
Ke & Hb much larger than Yg, Wf

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### Systematic delay

- Hb consistent from 2011-2016
- Ke most prominent in 2014
- Direct correction in analysis problematic:
- Rapid phase wrap
- Sometimes causes clock jumps
- → Correction model



Dan MacMillan calculated a correction model and applied it in the analysis

Unfortunately, no real improvement was found in the CRF

