

X/Ka Frame Improvements: Vision to Reality





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 ⁴⁸ 10 Feb 2010 - Session 04-T03

JPL/Caltech NASA CL#10-0948 10 Feb 2010, C.S. Jacobs



Outline



- Motivation for a Celestial Frame above 8 GHz
- X/Ka-band frame as it exist today
 - Accuracy: X/Ka vs. S/X based ICRF-2
 - Spatial Coverage
- Reducing the X/Ka error budget:
 - SNR: improved data rate 4X now, 40X within few years
 - Instrumentation:
 - Ka-band phase calibrator
 - Digital Back End
 - Troposphere calibration: (cheaper?) WVRs



Motivation



• Astrometry, Geodesy and Deep Space navigation, now at 8.4 GHz (X-band) with 2.3 GHz (S-band) plasma calibrations

Going to Higher radio frequencies allows

- More *compact* sources which should lead to more *stable* positions
- Higher Telemetry Rates to Spacecraft
- Smaller, lighter RF spacecraft systems
- Avoid S-band RFI issues
- Ionosphere & solar plasma down 15X !! at 32 GHz (Ka-band) compared to 8 GHz thus observe closer to Sun & Galactic center

Drawbacks of Higher radio frequencies:

- More weather sensitive, higher system temp.
- Shorter coherence times
- Weaker sources, Many sources resolved
- Antenna Pointing more difficult



Picture credit: SOHO/ESA/NASA





Results from Deep Space Network 8.4/ 32 GHz observations

The X/Ka Frame current status





ΔRA , $\Delta Dec: X/Ka vs. S/X ICRF2$

Accuracy tested vs. S/X ICRF2 the current IAU standard



RA: 190 µas = 0.9 nrad

Dec: 265 µas = 1.3 nrad

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X/Ka vs. S/X-based ICRF-2



323 common sources--at least 2 group delays, < 5mas sigma





X/Ka results: 351 Sources detected



Recent work fills holes: near Galactic plane & in south 17 candidates (marked '?') await correlation



Improving the X/Ka frame for the future



Attacking the X/Ka Error budget:

- SNR can be improved +8 dB!
- Instrumentation: Phase calibrators Digital Back End
- Troposphere cals: WVR





Results have been limited by Ka-band SNR

Solution:

1) More bits: 4X operational 8X R&D in ~6 months Will yield +5 dB SNR increase

2) Ka pointing

Now with improved Pointing calibrations ~3 dB more SNR

Total vs. early passes +8 dB SNR increase!



Results have been SNR limited for SNR < 15 dB



X/Ka delay precision



Deep Space Network 34m Beam-waveguide antennas

- 47 sessions, 10K delays/rates, median delay sigma 45 psec very small data set by S/X standards of 6 x 10e+6 delays
- Data rate: 43 passes @ 112 Mbps (X/Ka 56/56 Mbps)

3 passes @ 224 Mbps (X/Ka 80/144) ~ 3X

1 recent @ 448 Mbps (X/Ka 160/288) ~ 5X

in 6 mo. @ 896 Mbps (X/Ka 320/576) ~10X

Total Ka improvement 56 to 576 Mbps => 15 psec delay precision Reduces SNR below troposphere with +5.1dB Ka sensitivity!

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First Ka-band fringe @ 448 Mbps on Software Correlator





Credit: C. Jacobs, D. Bagri, E. Clark, C. Garcia-Miro, C. Goodhart, S. Lowe, E. Moll, L. Skjerve



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Results limited by No Ka-band Phase cal





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Credit: C. Jacobs, B. Tucker, L. Skjerve



Ka-band Phase Calibrator



• Concept: Tunnel diode Alan Rogers et al (Haystack)

• JPL prototype Ka-band phase cal: Hammel, Tucker, & Calhoun, JPL Progress Report, 2003

tmo.jpl.nasa.gov/progress_report/42-154/154H.pdf

• Production units: Blake Tucker



<u>Tunnel Diode Chip</u> 0.055" diameter by 0.020" thick Mounted on 0.119" diameter carrier for solid grounding



X/Ka Frame Improvements: Vision to Reality Ka-band phase calibrator





Direct interface to K connector inside coaxial structure.

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Credit: Blake Tucker

Pulse driver mounted as close as possible and fed through coaxial structure to minimize rise time and ringing



Ka-band phase calibrator





Tunnel diode module and pulse gating switch mount inside reference divider housing to form the comb generator module.



Oven control electronics stacks on top of the comb generator module.

Credit: Blake Tucker



Ka-band phase calibrator







Thermally controlled baseplate and housing salvaged from de-commissioned X-band comb generators.

Comb generator module and oven control electronics mounted inside oven.

Power conditioning mounted on outer cover to dissipate heat.

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Credit: Blake Tucker



Ka-band phase calibrator





Completed Ka-band phase calibrator with 32 GHz output signal. Credit: Blake Tucker



Improving the X/Ka frame for the future



Attacking the X/Ka Error budget:

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- Instrumentation: Phase calibrators
 Digital Back End
- Troposphere cals: WVR





Instrumentation: Digital Back End

• Digital Back End:

Replace 30 yr old analog Mk3/4 Data Acquisition Terminal with

- Digital phase linear FIR filters
- Programmable gate array design: powerful & flexible

> Spanned bandwidth increases from 360 to 500 MHz> Data rates increase from 1024 to 4096 Mbps max

Digital Back End team:

Robert Navarro - *lead (see poster paper)* Charles Goodhart Robert Proctor Steve Rogstad Elliott Sigman Melissa Soriano Duo Wang Les White



Digital Back End: hardware in the lab





IF select switch: 12 inputs allows multiple bands, multiple antennas



Command & Control

Mark-5C recorder





Sampler: 1280 MHz, 8-bit/sample

Copper to fiber, Digital filter, Format

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Design: Navarro et al. Photo credit: Les White





Summary of Instrumental Improvements

Instrument	MkIV	DBE/Mk5-C Comment
Filters	Analog 7-pole Butterworth	Digital FIR removes phase phase linear ripple in channel
Spanned bandwidth	360 MHz	500 MHz Mk4 baseband limit 1.4X improvement
Data rate @ start @ max.	112 Mbps 896 Mbps	2048 MbpsDSN SNR limited4096 Mbpstrop/inst. limited6X sensitivity
Phase Cal: S/X X/Ka	Yes No	Yes Yes removes 100s of psec



Improving the X/Ka frame for the future



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- Instrumentation: Phase calibrators Digital Back End

Troposphere cals: WVR





Troposphere Solution 2: Better Calibration

- JPL Advanced Water Vapor Radiometer

 1 deg beam better matches VLBI improved gain stability improved conversion of brightness temperature to path delay
- **Demos show 1mm accuracy!** Goldstone-Madrid 8000 km baseline using X/Ka phase delays (Jacobs *et al*, AAS Winter 2005). (Bar-Sever et al, IEEE, 2007).
- A-WVRs deployed at Goldstone & Madrid Seeking funding for Tidbinbilla, Australia
- A-WVR will not be used operationally for X/Ka catalog until SNR and phase cal errors are peeled back.



VLBI Delay Residuals DOY 200 Ka-Band DSS26-DSS55







Troposphere Solution 3: Cheaper WVR?

- Advanced R&D WVR is very accurate but expensive on order ~\$500K Mounted 50-100m from VLBI antenna
- Investigating "Radiometrics" WVR
 - 21 channels cover 20-30 GHz
 - "Inexpensive" at \$120K (U.S.)
 - Light weight: ~7 kg
- Subreflector mounting
 - enabled by low weight
 - ideal co-pointing, no offset
 - better match with VLBI beam volume
 - enables calibration of doppler/rate
 - improves performance at high freq.

• Gain stability needs investigation





Summary



• X/Ka Celestial Frame: 351 sources detected Observations: 10K observations from 47 sessions Deep Space Net 34m beam waveguide antennas 8,400 and 10,500 km baselines

> Accuracy: X/Ka vs. S/X ICRF2 common 323 sources ΔRA ~ 190 μas wRMS (0.9 nrad) ΔDec ~ 265 μas wRMS (1.3 nrad)

• Near term potential 100 µas accuracy

Attacking three main error sources:

SNR: +8 dB of SNR improvements almost done!

Increased bit rate: ~5dB increase in SNR (~6 month) Pointing cals: ~3dB increase in SNR (done) Instrumentation: Ka-band phase calibrators & Dig. Back End being built Troposphere: Estimation using spatial/temporal correlations Calibration using (cheaper?) Water Vapor Radiometers